# Presenter

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## Background

Solar flares are multi-scale phenomena. The overall structure evolves on slow timescales where the MHD approach can be used. Near magnetic reconnection sites, the timescales are much faster, and a kinetic approach, such as Particle-In-Cell (PIC), is needed. It is too computationally expensive to fully model a flare using only PIC.

We here present the preliminary validation results of a new PIC solver based on the Photon-Plasma<sup>[1]</sup> code. The solver is implemented in the DISPATCH<sup>[2]</sup> framework which allows dynamic switching of solvers in each 'patch







# UNIVERSITY OF OSLO **Towards Realistic Solar Flare Models** Particle-In-Cell simulations in the exascale era

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Each 'patch' in DISPATCH has a local timestep and may run different

With AMR different temporal and spatial scales can be run concurrently in the same simulation.

flares.

Vay Particle velocity and position.

References:

Michael Haahr<sup>1,2</sup>, Boris V. Gudiksen<sup>1,2</sup>, & Åke Nordlund<sup>1,3</sup>





T. Haugbølle, J. T. Frederiksen, Å. Nordlund, Photon-Plasma: A modern high-order particle-in-cell code, Physics of Plasmas 20 issue 6 (2013) [2] Å. Nordlund, J. Ramsey, A. Popovas, M. Kuffmeier, Dispatch: A numerical simulation framework in the exa-scale era - i. Fundamentals, Monthly Notices of the Royal Astronomical Society 477 (2018) [3] J. L. Vay, Simulation of beams or plasmas crossing at relativistic velocity, Physics of Plasmas 15 issue 5 (2008)



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